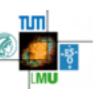


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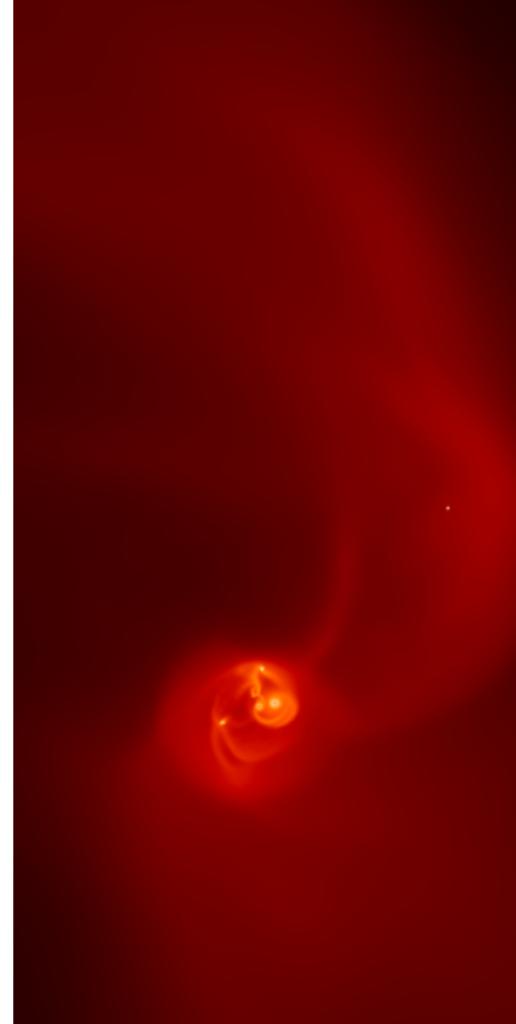


### Running Simulations with GANDALF

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#### The GANDALF parameters file

- The GANDALF parameters file is used to control almost all other aspects of the simulation, the generation of initial conditions and of the algorithms used.
- The parameters file has a simple structure :
- There are **way too many** parameters to go through each in detail, so we'll just go over the broad categories of parameters available and concentrate on a few important ones

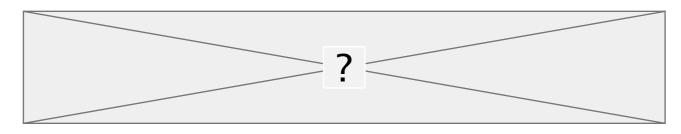
#### Core parameters

- ndim : Simulation dimensionality (1, 2 or 3)
- sim : Simulation type

sph	= SPH (+ N-body) algorithm (default : 'grad-h' SPH)
gradhsph	= 'grad-h' SPH simulation (+ N-body)
sm2012sph	= Saitoh & Makino (2012) SPH (+ N-body)
meshlessfv	= Meshless Finite-Volume algorithm (default : 'mfvmuscl')
mfvmuscl	= Meshless FV MUSCL integration simulation
mfvrk	= Meshless FV Runge-Kutta integration
nbody	= N-body only simulation

- nbody : Main N-body integration algorithm
  - lfkdk= 2nd-order Leapfrog kick-drift-kicklfdkd= 2nd-order Leapfrog drift-kick-drifthermite4= 4th-order Hermite schemehermite4ts= Time-symmetric 4th-order Hermite scheme

#### Core parameters



- in\_file\_form : Format of initial conditions file column = Simple column data format sf/seren\_form = SEREN ASCII format su/seren\_unform = SEREN binary format
- out\_file\_form : Format of outputted snapshot files column = Simple column data format sf/seren\_form = SEREN ASCII format su/seren\_unform = SEREN binary format
- tend : Termination time of the simulation (given in tunits)
- dt\_snap : Snapshot time interval (given in tunits)
- tsnapfirst : Time of first snapshot (given in tunits)

#### Scaling parameters

- dimensionless : Are all quantities dimensionless? (0 or 1)
- routunit : Position unit

pc/kpc/mpc	= parsec/kiloparsec/megaparsec
au	= astronomical unit
r_sun	= Solar radius
r_earth	= Earth radius
cm/m/km	= centimetre/metre/kilometre

• moutunit : Mass unit

m_sun	= Solar mass
m_jup/m_earth	= Jupiter mass/Earth mass
g/kg	= gram/kilogram

- toutunit : Time unit yr/myr/gyr = year/megayear/gigayear day = day sec = second
- voutunit : Velocity unit cm\_s/m\_s/km\_s = centimetres/metres/kilometres per second au\_yr = astronomical units per year

#### Hydrodynamical parameters

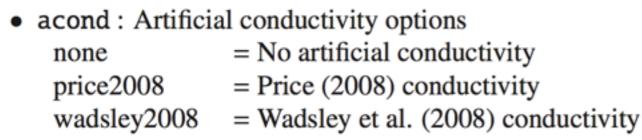
- hydro\_forces : Compute hydro forces? (1 or 0)
- gas\_eos : Gas particles equation-of-state
  - energy\_eqn = Solve energy equation
  - isothermal = Isothermal EOS

#### barotropic = Barotropic EOS (i.e. for mimicing isothermal + adiabatic phase during protostellar collapse)

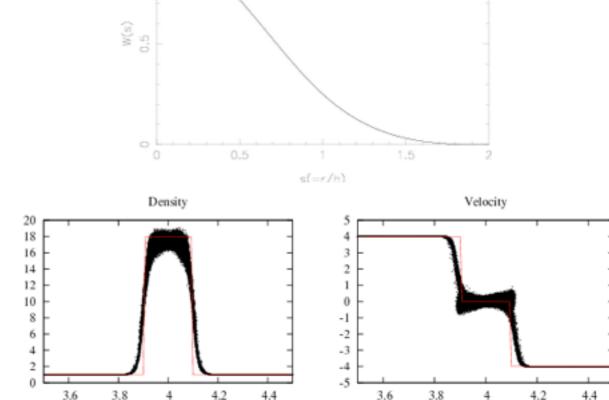
- barotropic2 = Similar to barotropic, but using discrete power laws rather than smooth change
- $rad_ws = EOS$  relating to Stamatellos et al. (2007) cooling method
- energy\_integration : Energy integration scheme (only applicable if solving the energy equation)
  - null = Energy equation not integrated separately
  - rad\_ws = Integrate energy terms using Stamatellos et al. (2007) method
- gamma\_eos : Ratio of specific heats for gas
- temp0 : (Isothermal) temperature (isothermal or barotropic EOS)
- mu\_bar : Mean gas particle mass (in units of hydrogen mass)

#### SPH parameters

- sph\_integration : SPH particle integration scheme
  - lfkdk = 2nd-order Leapfrog kick-drift-kick
  - lfdkd = 2nd-order Leapfrog drift-kick-drift
- kernel : SPH kernel function m4 = M4 Cubic spline kernel quintic = Quintic spline kernel gaussian = Gaussian kernel (truncated at 3h)
- avisc : Artificial viscosity options none = No artificial viscosity mon97 = Monaghan (1997) viscosity

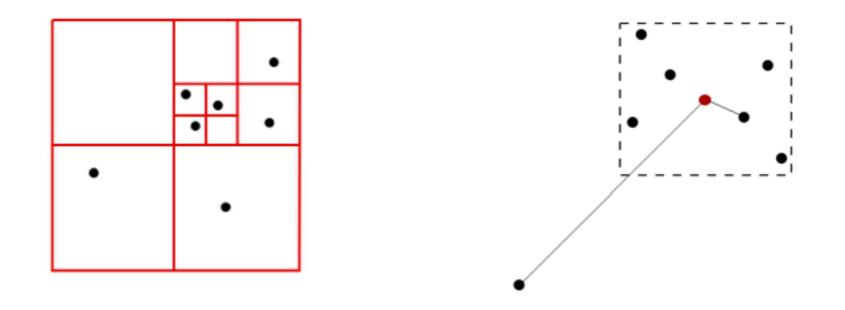


- time\_dependent\_avisc : Morris & Monaghan time-dependent viscosity (1 or 0)
- alpha\_visc : (Maximum) value of alpha viscosity parameter
- alpha\_visc\_min : Minimum value of alpha for time-dependent viscosity
- beta\_visc : Value of beta viscosity as a multiple of alpha



#### Tree parameters

- neib\_search : Neighbour searching algorithm
   bruteforce = Brute-force (i.e. summation over all particles)
   kdtree = Balanced kd-binary tree
   octtree = Barnes-Hut octal tree
- Nleafmax : Maximum no. of particles allowed in tree leaf cell
- ntreebuildstep : Integer steps inbetween tree re-builds
- ntreestock : Integer steps inbetween tree re-stocks

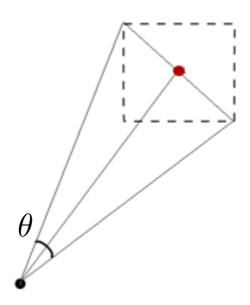


#### Gravity parameters

- self\_gravity : Compute gravitational forces? (1 or 0)
- gravity\_mac : Gravity-tree cell-opening criteria (N.B. always defulats to geometric for now)
  - geometric = Standard Barnes-Hut geomtric opening angle criterion
  - eigenmac = Compute eigenvalues of quadrupole moment tensor for MAC (Hubber et al. 2011)
- multipole : Multipole expansion for tree-gravity

monopole= Monopole-only terms for cell gravityquadrupole= Include quadrupole moment terms for cell gravityfast\_monopole= Compute monpoles more efficiently using Taylor expansion about cell COM

- thetamaxsqd : Maximum tree gravitational walk opening angle (squared)
- macerror : MAC error tolerance for individual cells
- external\_potential : External gravitational potential
  - none = No external potential
  - vertical = Constant gravitational field
  - plummer = Plummer background potential



#### Sink particle parameters

- sink\_particles : Do stars/sinks accrete? (0 or 1)
- create\_sinks : Create new sink particles? (0 or 1)
- smooth\_accretion : Use smooth accretion? (0 or 1)
- **rho\_sink** : Sink particle creation density (in cgs units)
- alpha\_ss : Sunyaev-Shakura alpha for smooth disc accretion
- **sink\_radius** : Sink particle radius (in units of smoothing length)
- sink\_radius\_mode : How to calculate new sink radius hmult = sink radius a multiple of SPH particle smoothing length fixed = sink radius is fixed for all new sinks

#### Initial conditions generators in GANDALF

- GANDALF contains a variety of in-built initial conditions generators for :
- 1D Hydrodynamical tests (e.g. shock-tubes, blast waves)
- Multi-dimensional hydrodynamical tests (e.g. Sedov-Taylor explosion, Kelvin-Helmholtz instability)
- Simple gravitational tests (e.g. free-fall collapse)

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- Simple N-body tests (e.g. binary stars, triple stars, Plummer sphere)
- Simplified astrophysical test cases (e.g. Boss-Bodenheimer test)
- Complete astrophysical initial conditions (e.g. turbulent prestellar core)
- Feel free to try other test problems (although only some of them you'll be able to plot with a simple plotting program like gnuplot)
- If you have splash successfully installed, then try changing the output format to 'sf' and then plotting them in splash

### Compiling the code

We wil compile GANDALF with the simplest possible set of options

CPP PYTHON COMPILER_MODE PRECISION OPENMP OUTPUT_LEVEL DEBUG_LEVEL	= DOUBLE = 0 = 1			
# FFTW libary flag	s and paths.			
"				
FFTW FFTW_INCLUDE	=			
FFTW_LIBRARY	=			
<pre># GNU Scientific library flags and paths. #</pre>				
GSL	= 0			
GSL_INCLUDE	=			
GSL_LIBRARY	=			

- To compile the full C++ code AND the python library :
- make —j

• To just compile the C++ executable :

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make –j executable

#### Running simulations on the command-line

- Once compiled, the gandalf executable will be placed in the 'bin' sub-directory located in the main gandalf directory :
- You can either :

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- Run it with the absolute path (e.g. bin/gandalf), or
- Set your PATH directory to include the gandalf bin subdirectory
- To run a simulation using the parameters file 'params.dat', type :

bin/gandalf params.dat

## Practical 1 : Run shocktube simulation in GANDALF

- Let's run some simple test problems with GANDALF
- From the 'tests' sub-directory, open the **adshock.dat** parameters file
- Run the simulation with

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bin/gandalf tests/adsod.dat

- The simulation should produce a series of output dumps of the form ADSHOCK1.su. 00001, ADSHOCK.su.00002, etc...OR ADSHOCK1.column.00001, ADSHOCK1.column.00002 (if you selected column format)
  - Plot the results with a simple plot program (e.g. gnuplot)

#### Practical 2 : Modify parameters in adsod.dat

Try experimenting with the parameters in the file, e.g.

- **double the output frequency** of snapshots
- **double the number of particles** in the simulation
- What happens if you **reduce or even switch off artificial viscosity**?
- change the SPH kernel

#### Restarting simulations

To restart a simulation using the last snapshot file generated, just run gandalf as usual but **with the '-r' option added**, i.e.

bin/gandalf -r params.dat

- The code produces a file called 'runid.restart' which contains the filename (and format) of the last snapshot produced by the code
- This could be used when :

- the simulation has crashed (or the computer has crashed)
- the simulation endtime has been reached and you wish to extend the simulation

#### Practical 3 : Restarting simulations

- Run the simulation and kill it before it reaches the end (N.B. you might need to increase the number of particles so it doesn't run too fast).
- Restart the simulation using the '-r' flag to verify it will successfully continue until the end
- Try changing the 'tend' parameter and restarting

### Practical 4 : Create a new simulation from a parameters file

**Create a new parameters file** (or copy an old one) to generate the following set of initial conditions for a shock problem :

- Isothermal equation of state, temp = 0.5
- LHS, rho = 1.0, vx = 0.0
- RHS, rho = 0.5, vx = -0.2
- tend = 0.2

- Choose appropriate values for other important parameters
- Plot the results at the end

# Compiling and running GANDALF for debugging

• GANDALF has a number of different options for helping with debugging

СРР	= g++
PYTHON	= python
COMPILER_MODE	= FAST
PRECISION	= DOUBLE
OPENMP	= 0
OUTPUT_LEVEL	= 1
DEBUG_LEVEL	= 0

- COMPILER\_MODE = DEBUG
  - Disables some optimisations and enables the '-g' flag (needed for debuggers)
- OUTPUT\_LEVEL = 2
  - Prints to screen more fine-grained information about where the code currently is running at
- DEBUG\_LEVEL = 1
  - Enables asserts in the code to help spot clear and quantifyable errors
  - $\mathsf{DEBUG\_LEVEL} = 2$

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• Enables more detailed (but very expensive) checking of individual algorithms

## Running GANDALF with gdb (and other debuggers)

- When COMPILER\_MODE = DEBUG, then you can run the code through the debugger (e.g. gdb, IIdb)
- To start the debugger (assuming gdb), type :

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gdb bin/gandalf

To run a simulation with a given parameters file, type :

run params.dat

- If the code crashes before the end, you can try various commands:
- to look at the subroutine call try to find out where the code crashed
- to print the values of the variables

### Quick reference list

- The debugger will stop automatically if you have a segmentation fault
- Once it's stopped you can **print** the value of local variables
- **bt** prints the full stack (i.e. tells you in which function you are, and which function called it, all the way up to main)
- list prints a few lines of code around the point where you stopped
- You can even execute your code one line at time using **step** (enters inside function calls) and **next** (stops once the function has finished). Useful when you want to see what the code is actually doing
- But true power comes with **breakpoints** -> tell gdb to stop at a particular line
- Use the following:

break file.cpp:15

and with **watchpoints** -> tell gdb to stop when a variable changes

watch variable
watch \*0x12345678

### Practical 5: Using the debugger 1

Type in the following code in the file 'debugtest.cpp'

```
#include <stdio.h>
void main()
{
    char *temp = "Paras";
    int i;
    i = 0;
    temp[3] = 'F';
    for (i=0; i<5; i++) printf("%c\n", temp[i]);
    return 0;
}</pre>
```

Compile with

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gcc –o –g debugtest debugtest.cpp

Now run with and without the debugger. See why a debugger is useful now?

### Practical 5 : Using the debugger 2

#### #include <iostream>

```
int ComputeFactorial(int number) {
    int fact = 0;
    for (int j = 1; j < number; j++) {
        fact = fact * j;
    }
    return fact;
}
int main() {
    int input;
    std::cout<< "Enter a number to compute its factorial" << std::endl;
    std::cin >> input;
    int fac = ComputeFactorial(input);
    std::cout << "The result is " << fac << std::endl;
}</pre>
```

- Locate and fix all problems with the debugger (even if your eagle-eye spots the problem by looking at the code) by stepping in the code line by line
- Remember to restart the debugger each time you re-compile