



Excellence Cluster Universe

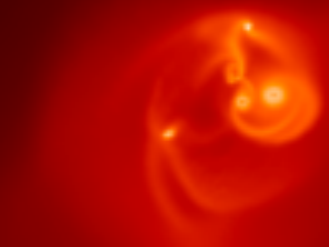


How to do units and scaling right (Or at least how to NOT do them wrong)

David Hubber

USM, LMU, München
Excellence Cluster Universe,
Garching bei München

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So, what's the wrong and right way to do units and why?

- Use physical units of the completely wrong scale (e.g. cm, s)

Very wrong!

- Use physical units of a more appropriate scale (e.g. pc, Myr)

Better but still wrong

- Use arbitrary dimensionless units (i.e. rely on the user to calculate any scale factors)

Correct, but more difficult than it needs to be

- Use dimensionless units where the scale factors are calculated by the code automatically

Yes!

What are dimensionless units exactly?

- **Dimensionless units** are a system of units where a physical quantity can be converted to an equivalent dimensionless quantity by way of a **scaling factor**, e.g.
- For example, take the quantity a which has a scale factor of A_0 , then we can convert this to an equivalent dimensionless quantity, a' :

$$a' = \frac{a}{A_0} \quad \Longleftrightarrow \quad a = A_0 a'$$

- Dimensionless units themselves **can always be converted back to the real, physical units** at any time
- They are different to **dimensionless constants** (such as e , π , etc..) which are truly dimensionless and are independent of any external system of units.

Why use dimensionless units?

Floating point precision

- One of the biggest arguments for using dimensionless units is the effect of **finite floating point precision**
- The single precision floating point range is $10^{-38} - 10^{+38}$
- Many astronomers love to **continue to use cgs units** (even in simulations) even though **they're completely inappropriate for almost any astrophysical context**
- Using cgs (or SI or similar non-astronomical units) can cause floating point precision in various situations, e.g.
- Computing volumes, e.g. a parsec size box volume,

$$1 \text{ pc}^3 = 2.93 \times 10^{49} \text{ m}^3 = 2.93 \times 10^{55} \text{ cm}^{-3}$$

- Computing quadrupole (or higher-order) correction terms in the gravity tree

$$\frac{1}{r^5} = \frac{1}{\text{pc}^5} = 3.57 \times 10^{-83} \text{ m}^{-5} = 3.57 \times 10^{-93} \text{ cm}^{-5}$$

Relations between dimensionless quantities

- Although we have freedom to choose a base set of dimensionless units (e.g. mass, length and time), derived quantities that use various combinations of these units must be consistent

$$r' = \frac{r}{R_0} \quad m' = \frac{m}{M_0} \quad t' = \frac{t}{T_0}$$

- Velocity units, for example, would be a combination of length and time units, i.e.

$$v' = \frac{v}{V_0} = v \frac{T_0}{R_0} \quad \text{where} \quad V_0 = \frac{R_0}{T_0}$$

- and acceleration units :

$$a' = \frac{a}{A_0} = a \frac{T_0^2}{R_0} \quad \text{where} \quad A_0 = \frac{R_0}{T_0^2}$$

Using dimensionless units to 'eliminate' physical constants

- Many interesting physical problems involves some physical constant, such as G , c , μ_0 , etc..
- We can chose a set of dimensionless units such that **the physical constant in the new set of units is unity**
- Various advantages to this
 - If chosen correctly, all physical quantities will be close to unity (better for summing floating point numbers, easier to spot 'incorrect' numbers)
 - Completely factors out needing to multiply by the constant (can save a little CPU time at least)

Example : Setting $G = 1$

- One common example of setting physical constants to unity is in gravitational problems. N-body codes would often employ a system of units that sets $G = 1$
- Substituting for our dimensionless units and rearranging

$$a = \frac{G m}{r^2} \quad \Rightarrow \quad \frac{R_0}{T_0^2} a' = \frac{G M_0 m'}{R_0^2 r'^2} \quad \Rightarrow \quad a' = \underbrace{\left\{ \frac{G M_0 T_0^2}{R_0^3} \right\}}_{G'} \frac{m'}{r'^2}$$

- The final equation in dimensionless form is similar with the constants grouped together, which we've called G' . If we wish to effectively set $G' = 1$, then this imposes a constraint on one of our quantities. Traditionally, this has been the time variable :

$$\frac{G M_0 T_0^2}{R_0^3} = 1 \quad \Rightarrow \quad T_0 = \left(\frac{R_0^3}{G M_0} \right)^{1/2}$$

- Typical units (in N-body and star formation problems) would select $R_0 = 1 \text{ pc}$ and $M_0 = 1 \text{ solar mass}$. What does the time unit, T_0 , come out as?

Example : Setting $G = 1$

$$T_0 = 4.7 \times 10^{14} \text{ s} = 14.91 \text{ Myr}$$

- This has a knock-on effect on any other unit that has time as a dimension
- e.g. Velocity unit

$$V_0 = \frac{R_0}{T_0} = 0.065 \text{ km s}^{-1} = 0.067 \text{ pc Myr}^{-1}$$

Another example : Setting $k_b/m_h = 1$

- In hydrodynamics, if we wish to convert from internal energy to temperature, we must use the Boltzmann constant, k_b , and the mass of a hydrogen atom, m_h .
- Similar to setting $G = 1$, we can set the collection of constants in the sound speed equation to unity to set an appropriate unit for temperature, i.e.

$$c^2 = \gamma \frac{k_b T}{\bar{m}} \quad \Rightarrow \quad V_0^2 c'^2 = \gamma \frac{k_b \theta_0 T'}{m_h \bar{\mu}} \quad \Rightarrow \quad c'^2 = \gamma \underbrace{\left\{ \frac{k_b \theta_0}{m_h V_0^2} \right\}}_1 \frac{T'}{\bar{\mu}}$$

$$\frac{k_b \theta_0}{m_h V_0^2} = 1 \quad \Rightarrow \quad \theta_0 = V_0^2 \frac{m_h}{k_b}$$

- Using the same typical star formation units we employed for the $G = 1$ example, we find that

$$\theta_0 = 0.51 \text{ K}$$

Units in GANDALF

- We created a 'SimUnit' class in GANDALF to hold and compute all required scaling factors
- Each required unit class inherits from this base class

```
class SimUnit
{
public:

    SimUnit();
    virtual ~SimUnit() {};

    virtual DOUBLE SIUnit(string) = 0;
    virtual string LatexLabel(string) = 0;
    DOUBLE OutputScale(string);

    DOUBLE inscale;
    DOUBLE inSI;
    DOUBLE outcgs;
    DOUBLE outscale;
    DOUBLE outSI;
    string inunit;
    string outunit;
};
```

```
class LengthUnit: public SimUnit
{
public:
    LengthUnit() : SimUnit() {};
    DOUBLE SIUnit(string);
    string LatexLabel(string);
};

class MassUnit: public SimUnit
{
public:
    MassUnit() : SimUnit() {};
    DOUBLE SIUnit(string);
    string LatexLabel(string);
};

etc..
```

Units in GANDALF

- An all-encompassing class called 'SimUnits' (there's an extra 's') which then holds everything in one place :

```
class SimUnits
{
public:

    SimUnits();
    ~SimUnits();

    void SetupUnits(Parameters *);
    void OutputScalingFactors(Parameters *);

    int dimensionless;           ///< Are we using dimensionless units?
    bool ReadInputUnits;        ///< Are input units read from snapshot?

    // Instances of all unit classes
    //-----
    LengthUnit r;               ///< Length unit
    MassUnit m;                 ///< Mass unit
    TimeUnit t;                 ///< Time unit
    VelocityUnit v;            ///< Velocity unit
    AccelerationUnit a;        ///< Acceleration unit
    DensityUnit rho;           ///< Density unit
    etc..
};
```

Input units vs Output units

- GANDALF is designed to handle simultaneously an input and output set of units
- For example, maybe you are reading in initial conditions in one set of units (e.g. pcs, Myr) but want to output in a different set (e.g. au, yr)
- However, in most cases, you will generate initial conditions with the same set of units OR create initial conditions on the fly
- Your choice of output units will often be set in the parameters file

```
#-----  
# Simulation units variables  
#-----  
Use physical units           : dimensionless = 0  
Length units                 : routunit = pc  
Mass units                   : moutunit = m_sun  
Time units                   : toutunit = myr  
Velocity units               : voutunit = km_s  
Density units                : rhooutunit = g_cm3  
Temperature units            : tempoutunit = K  
Specific internal energy units : uoutunit = J_kg  
Angular velocity unit        : angveloutunit = rad_s
```

Computing the scaling factors

- In order to simplify the calculation of scaling factors, **we calculate everything in the same set of units internally**. We use **SI units** (although we could have chosen cgs or another set if we wished)

Converts requested
output units to SI units

$$R_0 = R_{\text{outscale}} \times R_{\text{outSI}}$$

Converts code units to
requested output units

- If we have a different set of input and output units, then both of these should be consistent with each other, i.e.

$$R_0 = R_{\text{outscale}} \times R_{\text{outSI}} = R_{\text{inscale}} \times R_{\text{inSI}}$$

Length and mass units

- As discussed earlier, we select the length and mass units directly so these are trivial and are set as 1.0 each. We also need to compute the SI and cgs factors

```
// Length units
//-----
r.inunit   = params->stringparams["rinunit"];
r.outunit  = params->stringparams["routunit"];
r.inSI     = r.SIUnit(params->stringparams["rinunit"]);
r.outSI    = r.SIUnit(params->stringparams["routunit"]);
r.outcgs   = 100.0*r.outSI;
r.outscale = 1.0;
r.inscale  = r.outscale*r.outSI/r.inSI;
```

$$R_{\text{inscale}} = \frac{R_{\text{outscale}} \times R_{\text{outSI}}}{R_{\text{inSI}}}$$

Mass units

- As discussed earlier, we select the length and mass units directly so these are trivial and are set as 1.0 each

```
// Mass units
//-----
m.inunit   = params->stringparams["minunit"];
m.outunit  = params->stringparams["moutunit"];
m.inSI     = m.SIUnit(params->stringparams["minunit"]);
m.outSI    = m.SIUnit(params->stringparams["moutunit"]);
m.outcgs   = 1000.0*m.outSI;
m.outscale = 1.0;
m.inscale  = m.outscale*m.outSI/m.inSI;
```

Time units

- The time units are computed to ensure $G = 1$ as described earlier

```
// Time units
//-----
t.inunit   = params->stringparams["tinunit"];
t.outunit  = params->stringparams["toutunit"];
t.inSI     = t.SIUnit(params->stringparams["tinunit"]);
t.outSI    = t.SIUnit(params->stringparams["toutunit"]);
t.inscale  = pow(r.inscale*r.inSI,1.5)/sqrt(m.inscale*m.inSI*G_const);
t.inscale /= t.inSI;
t.outscale = pow(r.outscale*r.outSI,1.5)/sqrt(m.outscale*m.outSI*G_const);
t.outscale /= t.outSI;
t.outcgs   = t.outSI;
```

$$T_0 = \left(\frac{R_0^3}{G M_0} \right)^{1/2} \Rightarrow T_{\text{outscale}} \times T_{\text{outSI}} = \left(\frac{R_{\text{outscale}}^3 \times R_{\text{outSI}}^3}{G_{\text{SI}} M_{\text{outscale}} \times M_{\text{outSI}}} \right)^{1/2}$$

$$T_{\text{outscale}} = \left(\frac{R_{\text{outscale}}^3 \times R_{\text{outSI}}^3}{G_{\text{SI}} M_{\text{outscale}} \times M_{\text{outSI}}} \right)^{1/2} \frac{1}{T_{\text{outSI}}}$$

Converting initial conditions and input parameters to code units

- Converting initial conditions and/or parameters is simply a case of using the inscale/outscale variables.
- e.g. if you input velocities in km/s, then

```
Velocity units : voutunit = km_s
```

- Then all velocity quantities are scaled using $v' = \frac{v(\text{km_s})}{V_{\text{outscale}}}$
- If you have a parameter that is in either SI or cgs, then also divide by the SI/cgs factor

$$\rho' = \frac{\rho(\text{g cm}^{-3})}{\rho_{\text{outscale}} \times \rho_{\text{outcgs}}}$$

Code units vs output units

- GANDALF only uses these dimensionless units internally for computing quantities internally
- When generating output, such as snapshot files, all quantities are converted back to physical units (specified by the outscale selections)

$$v(\text{km}_s) = V_{\text{outscale}} \times v'$$

- For example, when outputting to file :

```
outfile << part.r[0]*simunits.r.outscale << "  "  
        << part.v[0]*simunits.v.outscale << "  "  
        << part.m*simunits.m.outscale << "  "  
        << part.h*simunits.r.outscale << "  "  
        << part.rho*simunits.rho.outscale << "  "  
        << part.u*simunits.u.outscale << "  "  
        << endl;
```

So, some basic rules about choosing units for your code

- Obviously, choose some set of units that is representative of your problem. Having a nice dimensionless framework means **nothing if you still chose strange units**
- **And that's about it!**